

Please ensure that your abstract fits into one column on one page and complies with the *Instructions to Authors* available from the Abstract Submission web page.

Do glasses in achondritic meteorites share a common source?

G. KURAT¹, M. E. VARELA², E. ZINNER³ AND
F. BRANDSTÄTTER¹

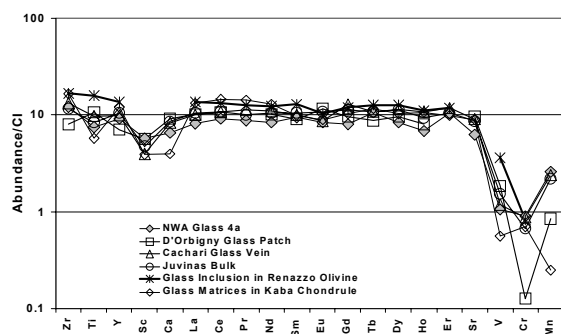
¹Naturhistorisches Museum, Postfach 417, A-1014, Vienna,

Austria, gero.kurat@univie.ac.at

²CONICET-UNS, Dpto. Geología, San Juan 670, 8000, Bahía Blanca, Argentina, evarela@criba.edu.ar

³Laboratory of Space Sciences and Physics Department, Washington University, St. Louis, MO 63130, USA, ekz@wuphys.wustl.edu

Howardites, eucrites and diogenites are breccia achondrites and are widely believed to be impactites from a parent body [1-3]. Howardites and eucrites also contain glass objects with major element chemical compositions that are similar to those of their respective bulk rocks – an apparent support of this belief. The recently discovered howardite NWA 1664 has abundant glass objects that are heterogeneous in their major and minor element abundances [4]. This suggests individual formation and processing of the glass objects rather than simple mixing and (shock) melting of anorthite+pyroxene. In contrast to the major elements, refractory lithophile trace element (TE) abundances in glass objects are unfractionated at 10 × CI (Fig.). They obviously signal a common homogeneous source that had chondritic relative TE abundances. Furthermore, abundances of TEs in glasses from NWA 1664 are similar to those in glasses from the eucrite Cachari and the angrite D'Orbigny and to the bulk eucrite Juvinas [5].



The very similar refractory TE abundances in glasses of howardites, angrites and eucrites likely indicate derivation from a common source. Achondrite glass objects share this feature with Ca,Al-rich glasses from glass inclusions in olivines of carbonaceous chondrites (Fig.) [6-8]. Consequently, all these glasses seem to have a common source with chondritic relative abundances of refractory TEs.

References: [1] Duke M. and Silver L. (1967) *GCA* 31, 1637-1665. [2] Bunch T. E. (1975) *Proc. Lunar Sci. Conf.* 6th, 469-492. [3] Mason B. (1967) *GCA* 31, 107-115. [4] Kurat G. et al. (2003) *LPSC* 34, #1832. [5] Varela M. E. et al. (2003) *GCA*, submitted. [6] Kurat G. et al. (1997) *MAPS* 32, Suppl., A76. [7] Varela M. E. et al. (2002) *MAPS* 37, Suppl., A142. [8] Varela M. E. et al. (2002) *GCA* 66, 1663-1679.